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ABSTRACT

The Southern Ocean's Antarctic Circumpolar Current (ACC) and Merid-16 ional Overturning Circulation (MOC) response to increasing zonal wind stress 17 is, for the first time, analyzed in a high- resolution (0.1° ocean and 0.25° atmo-18 sphere) fully-coupled global climate simulation using the Community Earth 19 System Model. Results from a 20-year wind perturbation experiment, where 20 the Southern Hemisphere zonal wind stress is increased by 50% south of 30° S, 2 show only marginal changes in the mean ACC transport through Drake Pas-22 sage; an increase of 6% (136 to 144 Sv) in the perturbation experiment com-23 pared with the control. However, the upper and lower circulation cells of the 24 MOC do change. The lower cell is more affected than the upper cell with a 25 maximum increase of 64% versus 39% respectively. Changes in the MOC are 26 directly linked to changes in water mass transformation from shifting surface 27 isopycnals and sea ice melt, giving rise to changes in surface buoyancy forc-28 ing. The increase in transport of the lower cell leads to upwelling of warm 29 and salty Circumpolar Deep Water and subsequent melting of sea ice sur-30 rounding Antarctica. The MOC is commonly supposed to be the sum of two 3 opposing components: a wind- and transient-eddy overturning cell. Here, the 32 transient-eddy overturning is virtually unchanged and consistent with a large-33 scale cancellation of localized regions of both enhancement and suppression 34 of eddy kinetic energy along the mean path of the ACC. However, decompos-35 ing the time-mean overturning into a time- and zonal-mean component and a 36 standing-eddy component, reveals partial compensation between wind-driven 37 and standing-eddy components of the circulation. 38

3

39 1. Introduction

It is currently estimated that more than 40% of the oceanic uptake of anthropogenic CO₂ takes 40 place south of 40°S (Sallée et al. 2012) and mesoscale eddies play an important role in the up-41 take (Gnanadesikan et al. 2015). The oceanic uptake over the Southern Ocean is largely governed 42 by the strength of the meridional overturning circulation (MOC) and the location of outcropping 43 isopycnals at the surface (Marshall and Speer 2012; Morrison et al. 2015). Over the past fifty 44 years, Southern Ocean winds have been increasing at a steady rate and have shifted poleward in 45 response to anthropogenic forcing from the Antarctic ozone hole in the lower stratosphere and 46 global climate change (Thompson et al. 2011). Recent work from paleo records suggests that the 47 Southern Ocean winds have been weaker in past climates due to an equatorward shift of the polar 48 Westerlies (Toggweiler 2009) and are currently the strongest they have been in the past thousand 49 years (Abram et al. 2014). Understanding how the strength of the Antarctic Circumpolar Cur-50 rent (ACC) and MOC respond to changing winds is fundamental to understanding global climate 51 change. 52

The ACC and MOC response to changing surface wind and buoyancy forcing hinges on the 53 response by mesoscale eddies. One hypothesis is that there will be zero change in ACC trans-54 port with increasing winds. The paradigm is that as winds increase, the ACC transport remains 55 roughly constant, because wind-forced steepening of isopycnals will quickly be brought back to 56 their original state. This is accomplished by down-gradient eddy buoyancy fluxes, generated via 57 baroclinic instability, that draw on the excess available potential energy inparted by the increase 58 in zonal wind stress. Interfacial form stress, which is proportional to the eddy buoyancy fluxes, 59 then transfers this excess wind-driven momentum to the sea floor where it is balanced by bottom 60 form drag (Ward and Hogg 2011). The near independence of the ACC transport (usually through 61

⁶² Drake Passage) to changes in the Southern Hemisphere wind stress is commonly referred to as ⁶³ *eddy saturation* (Straub 1993; Munday et al. 2013). Eddy saturation is becoming more widely ⁶⁴ accepted as limited observations (Firing et al. 2011; Chidichimo et al. 2014; Böning et al. 2008) ⁶⁵ and modeling efforts (Farneti et al. 2015) see only small trends in ACC transport over multiple ⁶⁶ decades of increased wind forcing.

It is less clear how the Southern Ocean's zonally-integrated meridional circulation responds to 67 changing winds. Theories suggest that the Southern Ocean MOC is the small residual of a near 68 cancelation of two opposing meridional circulation cells; a clockwise (looking west) wind-driven 69 circulation in the density-latitude plane (equatorward at the surface and poleward at depth) known 70 as the "Deacon Cell" and an eddy-driven cell of opposite sense (Johnson and Bryden 1989; Mar-71 shall and Radko 2003). The near independence of the MOC to changes in wind stress is referred to 72 as *eddy compensation* because any changes to the wind-driven Deacon Cell will be compensated 73 by the eddy-driven circulation. Crucially, any changes in the MOC must be consistent with surface 74 buoyancy modifications, and thus water mass transformation, related to changes in the wind field. 75 The degree of eddy compensation presently taking place in the Southern Ocean is unknown, 76 but there is observational evidence from satellite altimetry that surface eddy kinetic energy (EKE) 77 has increased in recent decades (Meredith and Hogg 2006; Hogg et al. 2015). Oceanic observa-78 tions are too sparse to make a direct diagnosis of the MOC so that there is no "true" estimate for 79 models to use as a benchmark. Estimates of the MOC from models that most realistically repre-80 sent the ocean are instead used as a "true" depiction of the MOC. Recent work by Farneti et al. 81 (2015) and Downes and Hogg (2013) show that there is considerable spread in the strength of 82 the MOC across coarse-resolution (1°) models forced with the Coordinated Ocean-ice Reference 83 Experiments (CORE-II) winds (Large and Yeager 2009) and amongst coupled climate models re-84 spectively. Climate models rely on an accurate parameterization of mesoscale eddies, which are 85

not routinely resolved in climate models (Gent and McWilliams 1990) (hereafter referred to as
 GM). Idealized studies show that the MOC response is more sensitive to model resolution than is
 the ACC transport (Stewart et al. 2014; Morrison and Hogg 2013).

To address how the MOC will respond to changing winds, a wide range of model experiments 89 have been performed from idealized eddy-resolving zonally-re-entrant channels to more realistic 90 coarse-resolution global circulation models. Gent (2015) provides an extensive review of the ef-91 fects of Southern Hemisphere wind changes on the MOC in ocean models. Idealized studies that 92 are eddy resolving show that there is partial eddy compensation, but results are sensitive to sur-93 face boundary conditions (Abernathey et al. 2011; Wolfe and Cessi 2010). Gent and Danabasoglu 94 (2011) (hereafter referred to as GD11) show that in a realistic coupled climate model, eddy com-95 pensation is achieved if the GM eddy coefficient is allowed to vary in space and time in response 96 to changes in stratification, which is consistent with idealized studies where EKE increases with 97 winds (Abernathey et al. 2011; Morrison and Hogg 2013). The highest resolution simulations to 98 date to investigate MOC changes due to wind forcing are a Southern hemispheric eddy-permitting 99 model $(1/6^{\circ})$ (Hallberg and Gnanadesikan 2006) and a global eddy-permitting $(1/4^{\circ})$ simulation 100 (Treguier et al. 2010). However, both models' are uncoupled to the atmosphere and representation 101 of Southern Ocean mesoscale eddies might be under-resolved, because the first baroclinic defor-102 mation radius, which is the characteristic spatial length scale of mesoscale eddies and ranges from 103 6-25 km in the Southern Ocean (Smith 2007), requires higher than $1/6^{\circ}$ resolution south of 30° S 104 (Hallberg 2013). 105

To overcome uncertainty regarding the parameterization and resolution of mesoscale eddies, this study examines the ACC transport and MOC response to changing winds in a new, highresolution version of the Community Earth System Model (CESM). The experiment is conducted with a fully coupled configuration of CESM with an eddy resolving (0.1°) ocean component and high-resolution (0.25°) atmosphere, which also includes sea ice and land components at the same
resolution as the ocean and atmosphere respectively. The increase in resolution compared to a
standard CESM simulation at coarse resolution (1° ocean and atmosphere) can be seen from snap
shots of sea surface temperature (SST) and sea surface height (SSH) in Fig. 1. When mesoscale
eddies are explicitly resolved, there is more filamentary structure and closed contours of SSH.

With this high-resolution version of CESM we test the ideas of eddy saturation and eddy com-115 pensation by running a twenty-year wind perturbation (WP) experiment in which the zonal wind 116 stress in the Southern Ocean south of 30° S is increased by 50%, the same as the PERT1 experi-117 ment in GD11. This is the highest resolution simulation brought to bear on this problem to date. 118 A challenge of interpreting our results is that a twenty-year perturbation experiment will not have 119 equilibrated in the deep ocean. Other modeling efforts at eddy-permitting resolution (0.25°) have 120 experienced considerable model drift (Treguier et al. 2010). The emphasis of this study is on the 121 transient response of the upper part of the the ACC. However, the deeper ACC comes into play 122 through the momentum budget. With this experiment being the first of its kind at this resolution, 123 it is expected that these caveats will be refined over time. Model drift will be discussed further in 124 the model description section. 125

One of the interesting results of this study is that standing eddies play the dominant role in the 126 response of the MOC to increasing winds compared to transient eddies. Standing eddies were 127 recognized as a major contributor to meridional fluxes in the Southern Ocean by early studies 128 (de Szoeke and Levine 1981; Treguier and McWilliams 1990; Wolff et al. 1991). However, popular 129 theoretical models of the ACC are framed in terms of a "streamwise average," which follows the 130 meanders of the time-mean current, effectively eliminating the standing component and leaving a 131 balance between a wind-driven Ekman component and a transient eddy component (Marshall et al. 132 1993; Marshall and Radko 2003; Nikurashin and Vallis 2012). Recent studies have re-emphasized 133

the importance of standing eddies for the time-mean meridional flux of heat in the Southern Ocean
(Volkov et al. 2010; Bryan et al. 2014) and in the response to wind perturbations (Dufour et al.
2012; Viebahn and Eden 2012; Zika et al. 2013b; Thompson and Garabato 2014; Abernathey and
Cessi 2014).

Finally, we show that differences in surface buoyancy flux are closely tied to changes in the 138 MOC. For example, increased upwelling of cold water can lead to cooler SSTs and consequently 139 a decrease in latent and sensible heat fluxes. As described in the theoretical model of Marshall and 140 Radko (2003), the net upwelling and subduction in and out of the surface layer must be balanced 141 by diabatic processes (air-sea fluxes and mixing) within the surface layer. In an idealized, ocean-142 only, eddy-resolving model, Abernathey et al. (2011) showed how wind perturbation experiments 143 could lead to different MOC changes depending on the details of the surface buoyancy boundary 144 conditions. Only a coupled climate model, such as the one examined here, can hope to represent 145 the full range of feedbacks which govern the changes in SST, surface buoyancy flux, and MOC in 146 response to a wind perturbation. In Sec. 6, we examine the connection between changes in surface 147 buoyancy flux and changes in overturning through the lens of water mass transformation. 148

The paper is organized as follows. In the section 2 the MOC and its constituents are defined. In section 3 the control and WP experiments using high-resolution CESM are described. Section 4 shows the results of the ACC transport and MOC response in the WP experiment compared to the control simulation. Section 5 discusses the importance of the standing component of the mean overturning circulation in the MOC balance. Section 6 makes the connection between water mass modification and changes in the MOC. Lastly, section 7 contains a discussion and the conclusions of our results.

156 2. Meridional Overturning Circulation

The Southern Ocean MOC in depth space, referred to here as the Eulerian-mean MOC, is calculated by integrating the meridional velocity, v, zonally and vertically,

$$\Psi(y,z) = \overline{\oint \int_{z}^{0} \upsilon \, dz' \, dx},\tag{1}$$

where x and z are the zonal and vertical coordinates respectively. The overbar indicates a time 159 average, $\overline{()} \equiv \frac{1}{\tau} \int_0^{\tau} () dt$, where τ is the averaging period. This representation of the MOC is 160 largely made up of the wind-driven Ekman circulation which is well known as the "Deacon Cell." 161 It was later realized by Döös and Webb (1994) that the Deacon Cell vanishes when the MOC is 162 calculated vertically in density rather than depth space. The overturning circulation calculated 163 in density space better represents water mass transport, which is largely along isopycnals in the 164 ACC interior and consistent with the net overturning resulting from two opposing mechanisms; the 165 wind-driven circulation of the Deacon Cell and the eddy- driven circulation. The MOC is defined 166 as 167

$$\psi_{\text{moc}}(y,\sigma) = \overline{\oint \int_{\sigma}^{\sigma_s} v h \, \mathrm{d}\sigma' \, \mathrm{d}x},\tag{2}$$

where $h(x, y, \sigma, t) \equiv -\partial \tilde{z}/\partial \sigma$ is the thickness of isopycnal layers, \tilde{z} is the depth of isopycnal surfaces, and σ_s is the density value at the surface. σ is potential density and in practice σ_2 (potential density referenced to 2000 m depth) is used to calculate the MOC since its vertical structure tends to have monotonic profiles and is a good representation of the MOC calculated in neutral density classes (Lee and Coward 2003). As mentioned above, ψ_{moc} can be thought of as consisting of two opposing cells between the wind-driven overturning circulation, $\overline{\psi}$, and transient eddy-induced overturning circulation, ψ^* ,

$$\psi_{\text{moc}}(y,\sigma) = \overline{\psi} + \psi^*. \tag{3}$$

¹⁷⁵ The time-mean overturning circulation in isopycnal coordinates is defined as

$$\overline{\psi}(y,\sigma) = \oint \int_{\overline{\sigma}}^{\overline{\sigma}_s} \overline{\upsilon}\overline{h} \,\mathrm{d}\sigma' \,\mathrm{d}x, \tag{4}$$

where $\overline{h}(x, y, \sigma)$ is the time mean thickness of isopycnal layers. The transient-eddy-induced overturning circulation is then given by the difference between the MOC and time mean streamfunctions as

$$\psi^*(y,\sigma) = \psi_{\text{moc}} - \overline{\psi} = \overline{\oint \int_{\sigma}^{\sigma_s} \upsilon' h' \, \mathrm{d}\sigma' \, \mathrm{d}x},\tag{5}$$

¹⁷⁹ where the prime indicates a deviation from the time mean.

180 **3. Model**

181 a. Control experiment

The model used in this study is a high-resolution version of the CESM (Hurrell et al. 2013), 182 a new generation climate system model that is the successor to the Community Climate System 183 Model version 4 (Gent et al. 2011). Details of the simulation examined are summarized below, 184 but for a more in depth description see Small et al. (2014). The model configuration includes 185 the Community Atmosphere Model version 5 (CAM5) with a spectral element dynamical core, 186 Community Ice Code version 4 (Hunke and Lipscomb 2008), Parallel Ocean Program version 2 187 (POP2), and Community Land Model version 4 (Lawrence et al. 2011). CAM5 was integrated 188 with a horizontal resolution of about 0.25 degrees (specifically the spectral element dynamical 189 core with 120 elements on each face of the cubed sphere, referred to as ne120) and 30 levels in the 190 vertical. 191

The POP2 model has a nominal grid spacing of 0.1 degrees (decreasing from 11 km at the Equator to 2.5 km at high latitudes) on a tripole grid with poles in North America and Asia. The configuration is similar to that used in McClean et al. (2011) and Kirtman et al. (2012), except that the

number of vertical levels was increased from 42 to 62, with more levels in the main thermocline. 195 The ocean communicated with the coupler, providing updated SST and surface currents and re-196 ceiving updated surface fluxes, every 6 hours and the atmosphere communicated every 10 minutes. 197 The coupler computes air-sea fluxes using the Large and Yeager (2009) surface layer scheme. The 198 land and sea ice models are on the same grids as the atmosphere and ocean models respectively. 199 POP2 has been shown to produce eddy covariances consistent with observations in the Pacific 200 (Bishop and Bryan 2013) and Southern Ocean (Lenn et al. 2011), and this high-resolution version 201 of CESM through spectral analysis produces mesoscale eddy covariances of SST and geostrophic 202 meridional velocity which are consistent with satellite observations (Abernathey and Wortham 203 2015). 204

Following a 15-year spin up, the model was run for 66 years. We will refer to the last 66 years with model year 1 being equivalent to aggregate simulation year 16. For the current work we will focus on model years 45–66, which are the years when the WP experiment was performed and is described below. This time period was chosen because surface fluxes and ACC transport through Drake Passage had reached equilibrium by model year 45 (Small et al. 2014).

210 b. Wind Perturbation Experiment

The WP experiment is conducted with the same methodology as the PERT1 experiment in GD11 where the Southern Hemisphere zonal wind stress is increased by 50%, but here using CESM rather than CCSM4. The WP experiment is just short of a 21-year simulation, starting from March of model year 45 of the control simulation to the end of model year 65. The WP experiment was conducted by multiplying the zonal wind stress forcing the ocean component by 1.5 south of 35°S, with this factor linearly reducing to 1 to the north between 35° and 25°S and to the south between 65° and 70°S. The maximum time and zonal mean Southern Hemisphere wind stress is 41% larger compared with the control for model years 56–66 with an increase of 0.083 N m⁻² from 0.197 N m⁻² in the control to 0.280 N m⁻² in WP (Fig. 2a), which is very close to the maximum wind stress values cited in GD11. The increased zonal wind stress was not used in the bulk formulae to calculate the atmosphere-to-ocean heat and freshwater fluxes, and the increased zonal stress is not felt directly by the atmosphere component.

One of the complications of interpreting a perturbation experiment is model drift. As mentioned 223 before, model drift complicated results in lower resolution studies (Treguier et al. 2010). The 224 shortness of our WP experiment was dictated by very high computational costs. As demonstrated 225 in Small et al. (2014) and in Fig. 2b, ACC transport through Drake Passage is in equilibrium in 226 the control, exhibiting mainly interannual variability. After a decade of the WP simulation surface 227 variables reached a new equilibrium, which is demonstrated in surface EKE in Fig. 4 and described 228 in the next section. However, deep variables have not reached a new equilibrium. Fig. 3 shows 229 the monthly-averaged and area-averaged deep ocean temperature for the Southern Ocean in the 230 control and WP. The control deep ocean temperature is in equilibrium, albeit with some decadal 231 variability, but the WP is not. However, it is not clear whether a perturbation of this kind will have 232 reached equilibrium after 1000 years. 233

4. Results

²³⁵ a. Antarctic Circumpolar Current Transport Response

The ACC transport through Drake Passage is shown in Fig. 2b. The mean and standard deviation for the control and WP are 136 ± 3 Sv and 144 ± 6 Sv, respectively. The WP time series has a trend over the first ten years of 23 Sv decade⁻¹ with a small negative trend over the last ten years of -6 Sv decade⁻¹. It is not immediately clear why there is a small negative trend over the remaining ten years. The overall WP mean transport only increases by 6% with a 41% increase in the zonal
wind stress compared to the control simulation. These results suggest that the ACC transport is
largely *eddy saturated*.

The area-average surface EKE south of 30°S increased immediately following the change in 243 wind stress and then continued to increase for a period of 10 years, before stabilizing at a value 244 28% higher than the control mean (Fig. 4a). The area-average surface EKE has a linear trend 245 over the first ten years of 26 cm² s⁻² decade⁻¹, but no trend in the last ten years. However, the 246 spatial distribution of the EKE trends over the first ten years have positive and negative values that 247 are $O(\pm 100)$ cm² s⁻² decade⁻¹ (Fig. 4b), resembling the mean EKE difference over the last ten 248 years of the simulation (Fig. 5). Fig. 5 shows that there are both regions of enhancement and 249 suppression of EKE along the mean path of the ACC with increased zonal wind stress. Outside of 250 the Malvinas and the Agulhas return flows, the positive trends are largely in regions where standing 251 meanders of the mean flow occur in the lee of major topographic features, as shown in Thompson 252 and Garabato (2014) and Abernathey and Cessi (2014). Both the ACC transport through Drake 253 Passage and the area-averaged EKE have a linear trend during the first then years of the WP (Fig. 254 2b). Once the area-average EKE reaches an asymptote of \sim 35 cm² s⁻² above the control mean, 255 the ACC transport adjusts to the elevated EKE. For this reason overturning streamfunctions are 256 calculated for the final 10 years of the WP experiment (model years 56–66) in the next section. 257

The small increase in ACC transport is also reflected in the differences in isoypycnal surfaces between simulations (Fig. 6). The differences are largely focused in the upper ocean; in particular the top 100 meters within the mixed layer (Fig. 6b). There are only small changes in the interior (Fig. 6a). The top 100 meters is characterized by an increase in surface density in WP as compared to the control (Fig. 11a,b). There is also lighter (or more buoyant) waters at depth north of 45°S. This represents a surface intensified steepening of isopycnal surfaces, which makes up the increase ²⁶⁴ in the ACC transport. On average the water is more dense with a skewed probability density ²⁶⁵ distribution (Fig. 6c).

²⁶⁶ b. Meridional Overturning Circulation Response

267 1) EULERIAN-MEAN MERIDIONAL OVERTURNING CIRCULATION

The mean overturning streamfunction is made up of the clockwise (looking west) wind-driven 268 circulation in depth- latitude space, the Deacon Cell, and a weaker counterclockwise circulating 269 lower cell (Fig. 7). As expected with the increased winds in the WP simulation, the Deacon Cell 270 is intensified. The peak values of the Deacon cell are 40.5 Sv and 57.1 Sv in the control and WP 271 respectively (Fig. 7 a,b). The latitude and depth of the maximum values of the Deacon Cell are 272 similar between simulations. The latitude and depth of the maximum value of the Deacon cell in 273 the control simulation is 51.8°S and 607 m respectively (Fig. 7a). In WP they are 51.6°S and 552 274 m respectively (Fig. 7b). The maximum increase in the Deacon Cell between the control and WP 275 is 17 Sv and occurs at 48.3°S and at 830 m depth. This is equivalent to a 41% increase in transport, 276 consistent with a linear relationship with the surface wind stress (Fig. 7c). 277

The lower cell is reduced slightly between the control and WP (Fig 7c). The maximum values of the lower cell circulation are -14.7 Sv and -13.5 Sv in the control and WP respectively. The latitude and depth of the maximum value of the lower cell in the control simulation is 36.9°S and 3752 m respectively (Fig 7a). In WP they are 36.6°S and 3752 m respectively (Fig. 7b).

282 2) MERIDIONAL OVERTURNING CIRCULATION

The overturning streamfunction in isopycnal coordinates, referred to as the MOC, is estimated in this section for the control and WP experiments. Comparing the diagnosed MOC (Eq. 2) against a schematic of the overturning [Fig. 18 in Farneti et al. (2015)] there are four distinct circulation cells, but there are only two cells that change substantially between our simulations which will be described (Fig. 8a–c). There is a clockwise-rotating (looking west) upper cell centered at 48.1°S and $36.2\sigma_2$ with a maximum of 19.8 Sv in the control (Fig. 8a). The upper cell in WP is centered at 48°S and $36.35\sigma_2$ with a maximum of 27.5 Sv (Fig. 8b). The upper cell increased by a maximum of 7.7 Sv, which is a 39% increase over the control (Fig. 8c).

²⁹¹ There is a counter-clockwise-rotating cell called the lower cell that has an enhanced poleward ²⁹² region with a maximum of -20 Sv at 63.5°S and 37.15 σ_2 in the control (Fig. 8a). In WP the lower ²⁹³ cell has a maximum of -32.7 Sv at 64°S and 37.15 σ_2 (Fig. 8b). The counter circulation of the ²⁹⁴ lower cell increased by a maximum of 12.6 Sv in WP compared with the control, which is a 64% ²⁹⁵ increase over the control (Fig. 8c). The lower cell equatorward of 50°S has no difference between ²⁹⁶ simulations (Fig. 8c).

One decomposition of the MOC is between the time-mean overturning (Eq. 4) and the transient-297 eddy-induced overturning (Eq. 5) as stated in Equation 3. The time-mean overturning streamfunc-298 tion (Fig. 8d–f) has a nearly identical structure of circulation cells to the MOC (Fig. 8a–c). The 299 upper cell has a clockwise circulation with a maximum of 22.3 Sv at 52.2°S and 36.45 σ_2 in the 300 control (Fig. 8d). In WP the mean upper cell has a maximum of 30.7 Sv at 48.3°S and 36.30 σ_2 301 (Fig. 8e). The difference in the maximum values of the upper cell is 8.4 Sv, which is a 38% 302 increase over the control simulation (Fig. 8f). The lower cell has as a maximum value of -19.6 303 Sv at 64.5°S and 37.15 σ_2 in the control (Fig. 8d). In WP it has a maximum value of -30.9 Sv at 304 65.2° S and $37.15\sigma_2$ (Fig. 8e). The lower cell increases by 11.3 Sv, which is 58% greater than the 305 control (Fig. 8f). 306

The transient-eddy-induced streamfunction has a different structure than the MOC or the mean overturning (Fig. 8 g–i). There are two counter circulating cells: an upper and a deep cell. The upper cell is the stronger of the two and has a maximum of -16.7 Sv at 43.7°S and $35.30\sigma_2$ for the control (Fig. 8 g). The upper cell in WP has a maximum of -20.5 Sv at 43.4° S and $35.30\sigma_2$ (Fig. 8h). The small differences between the two simulations are at most ± 5 Sv concentrated in the upper cell with virtually no change between the lower cells (Fig. 8i). Overall, the transient eddy component does not make a large contribution to the zonal-mean overturning, as seen also in Dufour et al. (2012). However, an analysis in streamwise coordinates (not performed here) would likely find a much greater role for transient eddies (Viebahn and Eden 2012; Abernathey and Cessi 2014).

5. Time-Mean Meridional Overturning Streamfunction

In the previous section it was shown that little or no eddy compensation is apparent when the MOC is decomposed into time-mean and transient components. Much of the change in the MOC is reflected in the time-mean MOC (Fig. 8c,f). This prompted further investigation into the timemean overturning streamfunction. The time-mean overturning circulation, $\overline{\psi}$, can be decomposed into a time and zonal mean streamfunction, $[\overline{\psi}]$, and a deviation from the zonal mean as the "standing" component ψ^{\dagger} ,

$$\overline{\psi} = [\overline{\psi}] + \psi^{\dagger}. \tag{6}$$

 $\overline{\psi}$ is defined as

$$[\overline{\Psi}](y,\sigma) = \oint \int_{[\overline{\sigma}]}^{[\overline{\sigma}_s]} [\overline{\upsilon}][\overline{h}] \,\mathrm{d}\sigma' \,\mathrm{d}x \tag{7}$$

where $[\]$ is a zonal average, $[\] \equiv \frac{1}{L(y)} \oint (\) dx$, and L(y) is the circumpolar length of a given latitude circle. The residual overturning streamfunction is finally written as the three-part balance

$$\psi_{\text{moc}} = [\overline{\psi}] + \psi^{\dagger} + \psi^*. \tag{8}$$

The time and zonal mean component is essentially the Eulerian-mean MOC (Fig. 7) and remapped to the time and zonal mean depth of σ_2 surfaces (Fig. 9d–f). The differences in cir³²⁹ culation cells then are equivalent to the values stated for the Deacon Cell and lower cell in the ³³⁰ subsection 4.b. This decomposition is more physically insightful, because $[\overline{\psi}]$ is directly propor-³³¹ tional to the zonal mean Ekman transport and is therefore clearly wind driven.

The standing component of the MOC is shown in Fig. 9g–i. Similar to the transient component it has two counter cells that are comparable in strength: an upper and lower cell. However, the standing component is stronger than the time and zonal mean component. The upper cell has a maximum of -39.9 Sv at 51.5°S and $35.8\sigma_2$ in the control (Fig. 9g). The upper cell in WP has a maximum of -54.7 Sv at 51.1°S and $35.9\sigma_2$ (Fig. 9h). The upper cell increased by 14.8 Sv, which is 37% over the control (Fig. 9i).

The lower cell of the standing component has a maximum of -39.1 Sv at 57.8°S and 37.1 σ_2 in the control (Fig. 9g). In WP the maximum is -61.8 Sv at 57.8°S and 37.12 σ_2 (Fig. 9h). There is a 22.7 Sv increase in the lower cell circulation, which is a 58% increase over the control (Fig. 9i).

A large degree of compensation exists between the wind-driven time- and zonal-mean cell and the standing component of the overturning circulation, even though the peaks in overturning do not coincide in density-latitude space. In both the control experiment and in response to the wind perturbation,

$$\delta \psi_{\rm moc} \approx \delta[\overline{\psi}] + \delta \psi^{\dagger}, \tag{9}$$

since the change in transient eddy overturning is negligible, $\delta \psi^* \approx 0$ (Fig. 8i). δ here means WP minus CNTL. It is interesting to note that the lower cell does not exist in the time and zonal mean component of the overturning (Fig. 9d,e,f). It is only through the standing component of the overturning that the lower cell exists in our simulations (Fig. 9g,h,i).

The time-mean kinetic energy, $\overline{K} = \frac{1}{2}(\overline{u}^2 + \overline{v}^2)$, is also dominated by the standing component, which reflects changes in the zonal distribution of kinetic energy. The time-mean kinetic energy is decomposed into a time- and zonal-mean component and a standing component,

$$[\overline{K}] = \underbrace{\frac{1}{2}([\overline{u}]^2 + [\overline{v}]^2)}_{Mean} + \underbrace{\frac{1}{2}([u^{\dagger 2}] + [v^{\dagger 2}])}_{Standing}.$$
(10)

The biggest changes in the kinetic energy are in the standing component (Fig. 10). The longitudinal changes in the velocity field enhances mean kinetic energy as baroclinic eddy growth rates respond to changes in mean stratification and vertical shear (Thompson and Garabato 2014).

6. Surface Water Mass Transformation

Analysis of changes in the surface water mass transformation is used to better understand the 356 thermodynamics of the changes in the MOC. Here "water mass transformation" refers to the ther-357 modynamic modification of water density due to diabatic processes such as surface buoyancy 358 fluxes and mixing. When the transformation rates are integrated over isopycnals in an ocean 359 basin, the net transformation must balance the inflow / outflow in density coordinates, i.e. the 360 MOC defined in Eq. 2 (Walin 1982; Tziperman 1986; Speer and Tziperman 1992; Marshall et al. 361 1999; Marsh et al. 2000; Large and Nurser 2001; Iudicone et al. 2008; Downes et al. 2011). We 362 can therefore expect the changes in Eq. 2 under wind perturbation to be accompanied by changes 363 in water mass transformation rates. 364

³⁶⁵ With the wind perturbation there is an indirect effect on the coupled system through changes to ³⁶⁶ the surface properties, and this is reflected in changes to the net surface buoyancy flux (\mathscr{R}). Mean ³⁶⁷ changes in density, SST, and sea-surface salinity (SSS) are shown in Fig. 11. Surface density ³⁶⁸ increases almost uniformly along the ACC, with smaller patches of lighter waters found north of ³⁶⁹ 45°S (Fig. 11a). The zonally-averaged density distribution shows that the surface waters are more ³⁷⁰ dense at all latitudes except near 30°S and the increase peaks near 50°S (Fig. 11b). The changes in ³⁷¹ SST and SSS reflect this increase in density. Water spanning 10–20° latitude surrounding Antarctica are warmer and saltier by as much as $1-2^{\circ}$ C and 0.25-0.5 g kg⁻¹ respectively (Fig.s 11c,e). The warmer and saltier waters around Antarctica are mostly density compensated, but density does increase slightly. Water is cooler throughout most of the Pacific and Indian oceans, but salinity changes are less prominent with the exception that water is fresher near 30°S in the Pacific. The zonally-averaged SST and SSS show that the surface waters are more salty at latitudes south of 30° (Fig. 11f), but zonal average SST is only warmer south of 50°S (Fig. 11d).

³⁷⁸ Changes in the surface density field reflect changes to the surface buoyancy flux between the ³⁷⁹ simulations shown in Fig. 12. Surface buoyancy flux is the sum of surface heat flux (Q_o) and ³⁸⁰ fresh-water flux (FWF),

$$\mathscr{B} = \frac{\alpha_{\theta}g}{\rho_o c_p} Q_o - \frac{\alpha_S g}{\rho_{fw}} S_o(E - P - R)$$
(11)

where $\alpha_{\theta} = \partial \sigma / \partial \theta$, θ is potential temperature, $\alpha_S = \partial \sigma / \partial S$ and S is salinity, E - P - R is the 381 surface FWF (evaporation minus precipitation minus runoff), g is the acceleration due to gravity, 382 $\rho_o = 1026$ is the ocean reference density, $c_p = 3996$ J kg⁻¹ K⁻¹ is the specific heat at constant 383 pressure for seawater, $S_{\rho} = 34.7$ g kg⁻¹ is the ocean reference salinity for the virtual salt flux, and 384 $\rho_{fw} = 1000 \text{ kg m}^{-3}$ is the density of fresh water. The change in \mathscr{B} shows a complex spatial pattern 385 with a net buoyancy flux reduction near the coast of Antarctica and along the southern fringes of 386 the subtropical gyres. Along the core of the ACC, patches of both increased and decreased surface 387 buoyancy flux occur (Fig. 12a). The sign convention is such that of negative buoyancy flux means 388 the ocean is either cooling or becoming saltier. The zonally-averaged \mathcal{B} has sinusoidal meridional 389 structure (Fig. 12b). The changes in surface heat flux (Fig. 12c) dominate changes in surface 390 buoyancy forcing compared with FWFs (Fig. 12e). The FWF spatial difference (Fig. 12e) and 391 zonal average difference (Fig. 12f) both show positive buoyancy forcing near Antarctica. This is 392 consistent with a reduction in sea ice in the WP simulation (Fig. 13). In austral summer (JFM) and 393 winter (JAS) there is an overwhelming reduction in sea ice thickness surrounding Antarctica in the 394

WP simulation. Changes in FWFs around Antarctica though are masked by increased surface heat
 fluxes.

Since the MOC was analyzed in σ_2 coordinates, we use the same coordinate for the water mass analysis. To calculate the time-mean surface water mass transformation, the surface buoyancy flux (Eq. 11) is integrated over the surface outcrop area as

$$\Omega(\sigma_2) = \frac{1}{g} \frac{\partial}{\partial \sigma_2} \overline{\int_{\mathscr{A}_{\sigma_2}} \mathscr{B} \mathrm{d}A}$$
(12)

where \mathscr{A}_{σ_2} represents integration over the area south of 30° S with density greater than σ_2 . By breaking \mathscr{B} into Q_o and FWF components, Ω can be decomposed into transformation due to Q_o (Ω_{HF}) and FWF (Ω_{FWF}). There are additional contributions to water mass transformation due to interior mixing and cabbeling, but those are not diagnosed here (the necessary model output was not saved). As shown below, changes in surface transformation can explain most of the changes in the MOC.

Fig. 14 shows the annual mean Ω , Ω_{HF} , and Ω_{FWF} for both control and perturbation experi-406 ments. Examining the control case first, we see that Ω contains three peaks, each corresponding 407 with one of the cells of ψ_{moc} described above and shown in Fig. 8. (The transformation rates 408 should be compared with the MOC at 30° S.) For the densest waters (36.6 < σ_2 < 37.2 kg m⁻³) 409 surface cooling make water denser, with a peak transformation rate of 14 Sv; this corresponds 410 with the lower cell of the MOC. For water of intermediate density ($35.4 < \sigma_2 < 36.6 \text{ kg m}^{-3}$), 411 a combination of heat and FWFs (the dominant component) makes the water lighter, with a peak 412 transformation rate of -32 Sv; this corresponds with the upper cell of the MOC. Ω and ψ_{moc} do 413 not match up perfectly, since we have not calculated the transformation due to mixing; however, 414 this component can be inferred as the residual, as shown in Newsom et al. (2015). Mixing causes 415

additional water to be entrained into the subpolar cell; in contrast, mixing weakens the upper cell
 slightly and redistributes its position in density space.

The WP experiment produced a strengthening of both the upper and subpolar MOC cells. Corre-418 spondingly, in Fig. 14 we see that the transformation rates associated with these cells also increase 419 in magnitude. The increased heat loss associated with the lower cell causes transformation to 420 nearly double for water denser than $\sigma_2 = 37.0 \text{ kg m}^{-3}$, to a maximum of 26 Sv, matching the 12 421 Sv increase in overturning almost exactly. Although the upper cell transformation is dominated 422 by FWFs in the control, it is the heat fluxes which change most strongly under the WP; increased 423 density gain due to surface heat flux strengthens the upper-cell transformation to nearly -40 Sv and 424 shifts its maximum from $\sigma_2 = 36.0$ kg m⁻³ to $\sigma_2 = 36.2$ kg m⁻³, again consistent with the 8 Sv 425 increase in the upper cell. Changes in transformation due to FWF changes were minimal, as also 426 found by Newsom et al. (2015) in a greenhouse-warming scenario. 427

The changes in transformation are driven by the changes in heat flux shown in Fig. 12, namely 428 increased heat flux into the ocean in the ACC latitudes and increased heat flux out of the ocean 429 near Antarctica. The physical explanation for this change is consistent with increased Ekman 430 upwelling. In the ACC latitudes, where the water column is stably stratified in temperature, this 431 brings cooler water to the surface and, due to the interactive nature of latent and sensible heat 432 flux, produces increased heat gain. In the subantarctic region, where warmer water lies below 433 the surface, increased upwelling has the opposite effect, producing increased heat loss. Overall, 434 the entire pattern of Southern Ocean circulation and water mass transformation strengthens with 435 increasing winds. 436

437 7. Discussion and Conclusions

The Southern Ocean ACC transport and MOC response to changes in wind forcing in a fully-438 coupled high-resolution climate model (CESM) are diagnosed. Results from a \sim 21-year wind 439 perturbation experiment, where the Southern Hemisphere winds were increased by 50%, show 440 that the ACC transport is nearly eddy saturated, but the MOC is not eddy compensated. The 441 ACC transport response through Drake Passage only changes marginally. During the first ten 442 years of the simulation, while the eddies are ramping up, the ACC transport has a linear trend 443 of 23 Sv decade⁻¹ in response to the increased wind forcing. After a decade the eddies have 444 reached an equilibrium state and the ACC transport adjusted to the elevated level of EKE. The 445 overall ACC transport only increased by 6% compared to the control simulation suggesting that 446 the ACC is nearly eddy saturated. Recent studies have pointed to coastal winds near Antartica 447 as possible drivers of ACC transport variability (Zika et al. 2013a; Langlais et al. 2015), but the 448 time- and zonally-averaged wind in the WP experiment is indistinguishable from the control near 449 the Antarctica coast south of 65° S (Fig. 2). Thus, these potential mechanisms for ACC transport 450 variability are not explored in this paper. 451

The MOC increased in the upper and lower cells by 63% and 39% respectively. When the MOC 452 is decomposed into time-mean and transient-eddy contributions, the eddy-driven overturning does 453 not change much compared to the control. This result is consistent with the mean upper ocean 454 EKE difference (Fig. 5), where EKE is enhanced and suppressed along the mean path of the 455 ACC. The zonally-integrated mean EKE difference sums to near zero, which may help to explain 456 why our transient overturning difference is negligible. EKE is enhanced at choke points along the 457 mean ACC path near major topography and in the Western Boundary Current regions without any 458 systematic shifts in the Southern Ocean SSH fronts. 459

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Comparing the MOC results of this study with the same WP experiment (PERT1) performed 460 in the coarse-resolution CCSM4 climate model in GD11, we find a lesser role for transient ed-461 dies. The biggest changes in the overturning circulation in our high-resolution experiment arises 462 from enhancement of the steady rather than the transient eddy-induced overturning. When the 463 time-mean overturning circulation is decomposed into a time- and zonal-mean and a standing 464 component, the balance of the MOC is approximately between these two components (Eq. 9). As 465 the winds increase, the standing component of the overturning acts to partially, but not perfectly, 466 compensate for the increase in the wind-driven Deacon Cell. We now see a potential role for an-467 other type of eddy, the *standing eddy*, at higher resolution than Hallberg and Gnanadesikan (2006), 468 that plays the dominant role in poleward flow rather than transient eddies (Dufour et al. 2012) to 469 counter the wind-driven equatorward circulation. 470

In a coupled model changes in the MOC are connected directly with changes in surface wa-471 ter mass transformation. This is due to changes in the outcropping position of potential density 472 surfaces and the associated change in surface buoyancy forcing. The water mass analysis is a 473 diagnostic tool showing consistency between the two analysis methods. It is fundamentally prob-474 lematic to separate wind and buoyancy forcing in a coupled climate model to distinguish relative 475 contributions to changes in the MOC and ACC transport. Ultimately, in this model, the wind is re-476 sponsible for the changes in the MOC and ACC transport since this is the perturbed variable. Our 477 results suggest a key role for the ocean since changes to SST distributions (Fig. 11c) are accom-478 panied by changes to surface heating and cooling (Fig. 12c) that dominate the surface buoyancy 479 flux (Fig. 12a). 480

In this study we have simulated a \sim 21-year perturbation to the Southern Ocean that resolves mesoscale eddies. This simulation pushed the limits of the community's current computational capabilities. Ideally, analyzing the full response of the overturning circulation to a perturbation of

this type would require a simulation that is at least an order of magnitude longer. Thus, parame-484 terization of mesoscale processes is still needed. This simulation, however, does support the study 485 of Thompson and Garabato (2014), which suggests that the equilibration of the ACC to changes 486 in surface wind forcing will principally involve processes that are zonally-asymmetric. A change 487 in the wind stress without an increase in the zonal mean transport requires a greater vertical mo-488 mentum flux carried out by transient eddies. Thompson and Garabato (2014) argue that changes 489 in eddy activity will be focused in standing meanders related to fluctuations in the amplitude and 490 wavelength of the meander. It is important to note that this response is quite different from the 491 coarse-resolution study of GD11, where the GM coefficient increases almost uniformly through-492 out the ACC. A key result of this study is that directly resolving mesoscale eddies shows that the 493 response of ACC's EKE, for instance, to changes in wind stress is likely too nuanced to be de-494 scribed by trends spanning the whole ACC or even basins. We plan to analyze the momentum and 495 vorticity budgets of both high and low resolution simulations, similar to Cronin and Watts (1996) 496 and Hughes (2005) in a future study. 497

There are some potential shortcomings of our simulation. One shortcoming is the shortness of 498 the WP experiment. It is possible that there is decadal variability, but this is not possible to assess 499 in a \sim 21-year record. Another shortcoming is that the analysis was done using monthly archived 500 data. Ballarotta et al. (2013) found close correspondence in the MOC at 5-day versus monthly 501 archived data, suggesting that most of the signal is at monthly time scales and longer. However, 502 their analysis was done in an eddy-permitting model (0.25°) , which does not adequately resolve 503 the first baroclinic deformation radius at these latitudes (Hallberg 2013). Encouragingly, Aber-504 nathey and Wortham (2015), analyzing the same class of CESM simulation used here, found that 505 mesoscale eddy fluxes were dominated by sub-monthly frequencies. Future studies, for example, 506

could focus on longer simulations, meridional shifts of the mean winds to simulate past climates,
 and increases in vertical resolution.

With these caveats set aside, the high-resolution simulation has shown how increasing winds 509 could lead to an increase in the MOC over the time scales of twenty years. The results from 510 this study broadly agree with ozone depletion experiments, meant to simulate the increase in zonal 511 wind stress in the Southern Hemisphere during austral summer (Ferreira et al. 2015; Solomon et al. 512 2015). The average SST difference during the last ten years of our WP experiment resemble the 513 slow time-scale response in Ferreira et al. (2015), in which there is warming around Antarctica and 514 cooling at mid-latitudes (Fig. 11c). The wind stress changes are stronger in our WP experiment 515 than in the ozone depletion studies and as a result we see much more warming, up to 2° C, in 516 places around Antarctica. These results are indicative of an increase in the lower cell of the MOC 517 (Fig. 8), which causes more upwelling of warm and salty Circumpolar Deep Water (CDW). The 518 upwelling of CDW melts more sea ice throughout the year, which leads to an enhancement of 519 FWFs surrounding Antarctrica. The recent trends in winds observed over the past few decades in 520 Southern Hemisphere winds may already be having an influence on the MOC. This would infer 521 that the oceanic uptake of CO_2 may be changing as well. 522

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